



# STEREO Beacon

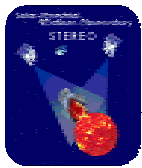


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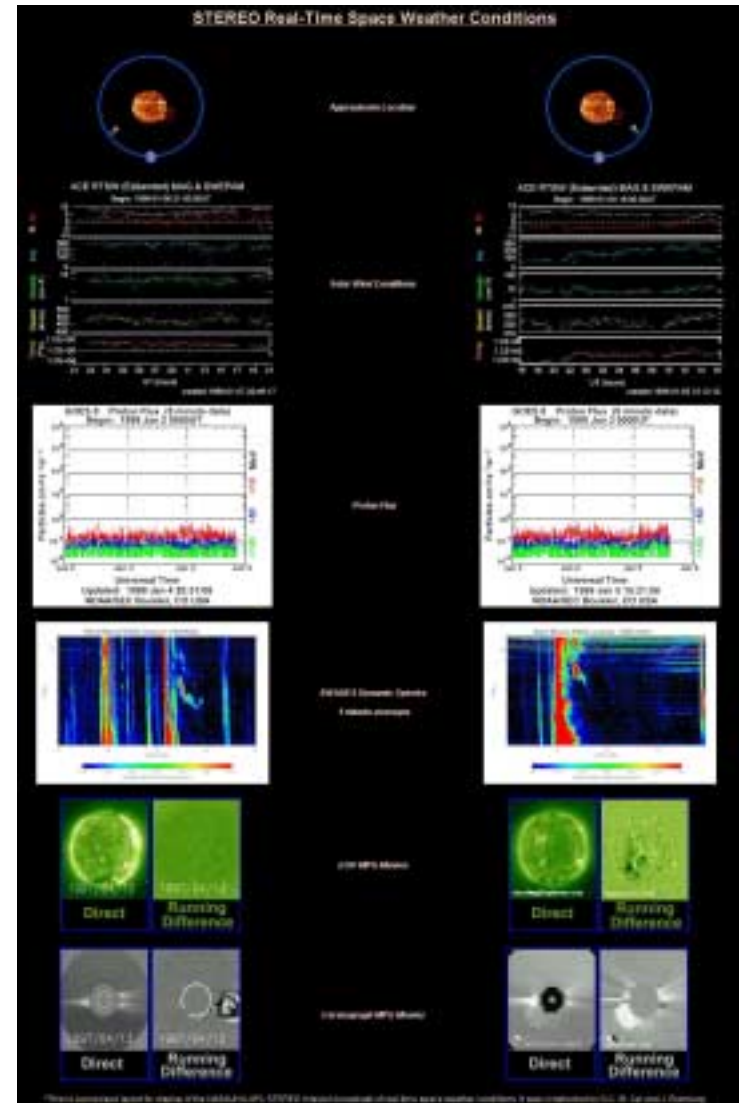




# STEREO Beacon



- Each STEREO spacecraft will broadcast highly compressed SWx images and *in situ* data continuously.
- During the daily 3-hour DSN contacts, the SWx stream will be captured, processed, and put online in near-real-time at the STEREO Science Center (SSC) located at NASA-GSFC.
- As soon as the recorder dumps are available, the SSC will fill in the previous 24 hours' SWx data as a browse archive.

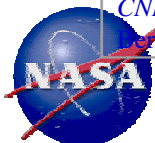


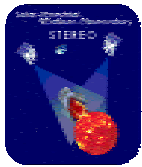


# STEREO Payload



<b>Instrument Name and Collaborating Institutions</b>	<b>Primary Measurement Space Weather Beacon Content [Number of 272 byte packets per minute]</b>
<p><b><u>IMPACT</u></b> (<i>In situ</i> Measurement of Particles and CME Transients)</p> <p><i>Principal Investigator: Dr. J. G. Luhmann, University of California, Berkeley, NASA-GSFC, Caltech, U. Md, U. Kiel, CESR, MPAe, JPL, ESTEC, UCLA, NOAA, LANL, et al.</i></p>	<p>Solar wind plasma characteristics; magnetic field parameters; solar energetic particles</p> <p>One minute average solar wind electron fluxes (6 energy bands); magnetic field strength and direction; energetic electron, proton, ion (He, CNO, Fe) fluxes (multiple bands) [1]</p>
<p><b><u>PLASTIC</u></b> (<b>PLAsma and SupraThermal Ion and Composition</b>)</p> <p><i>Principal Investigator: Dr. A. B. Galvin University of New Hampshire University of Bern, MPE-Garching, et al.</i></p>	<p>Ions in the energy-per-charge range of 0.2 to 100 keV/e</p> <p>One minute average solar wind proton density, bulk speed, thermal speed, and direction; alpha density; representative charge (or abundance) state distributions; suprathermal rates [1]</p>
<p><b><u>SECCHI</u></b> (<b>Sun-Earth Connection Coronal and Heliospheric Investigation</b>)</p> <p><i>Principal Investigator: Dr. R. A. Howard Naval Research Laboratory, Washington, D.C. Lockheed-Martin Solar and Astrophysics., NASA-GSFC, University of Birmingham (U.K.), IAS, RAL, MPAe, U. Kiel, CSL, et al.</i></p>	<p>EUV imager, two coronagraphs with overlapping fields of view; two heliospheric imagers with overlapping fields of view</p> <p>256x256 pixel highly compressed images from EUVI, COR1, COR2, HI1, HI2 [14]</p>
<p><b><u>S/WAVES</u></b> (STEREO/WAVES)</p> <p><i>Principal Investigator: Dr. J.-L. Bougeret CNRS, Observatoire de Paris, University of Minnesota, UC Berkeley, NASA-GSFC</i></p>	<p>Interplanetary radio bursts from 40 kHz to 16 MHz</p> <p>One minute average radio dynamic spectrum (Intensity, frequency, time) [1]</p>

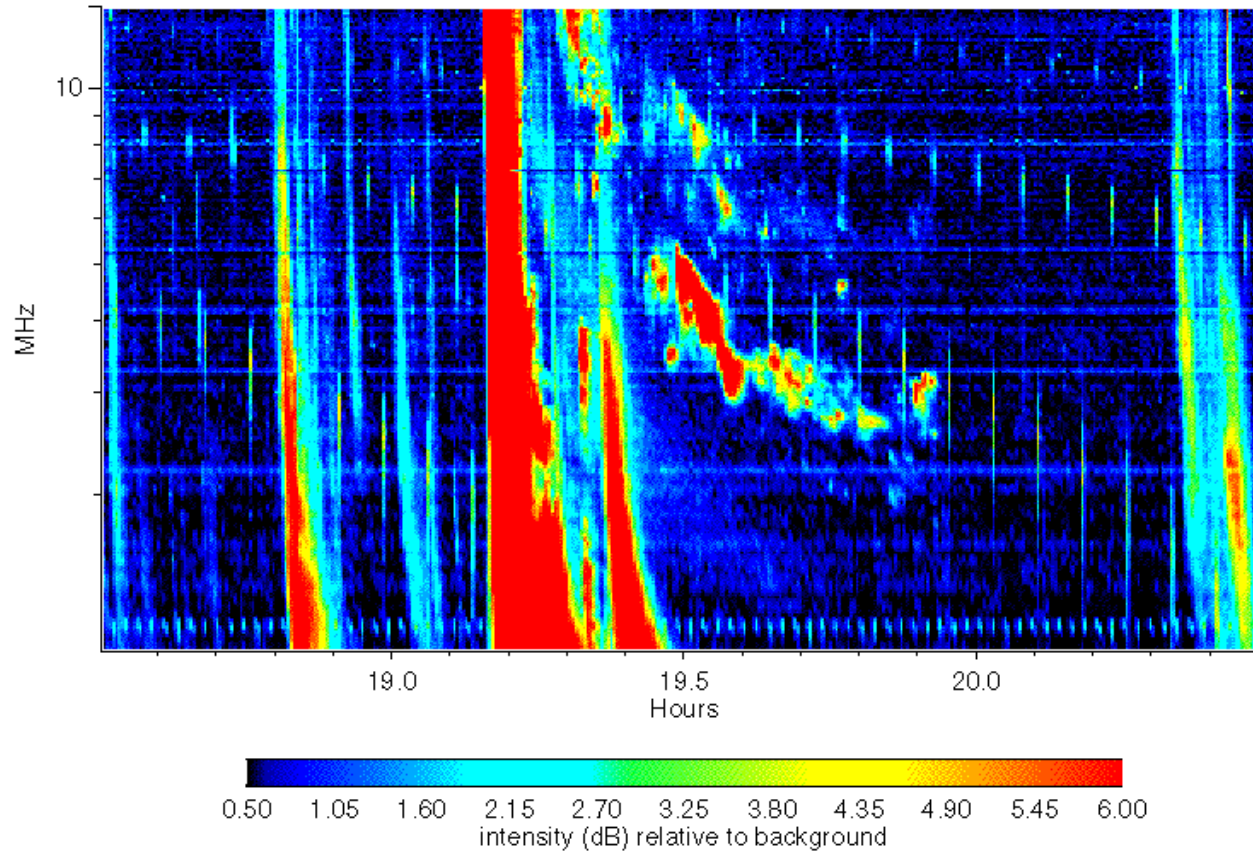


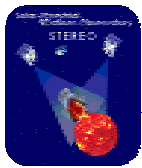


# S/WAVES



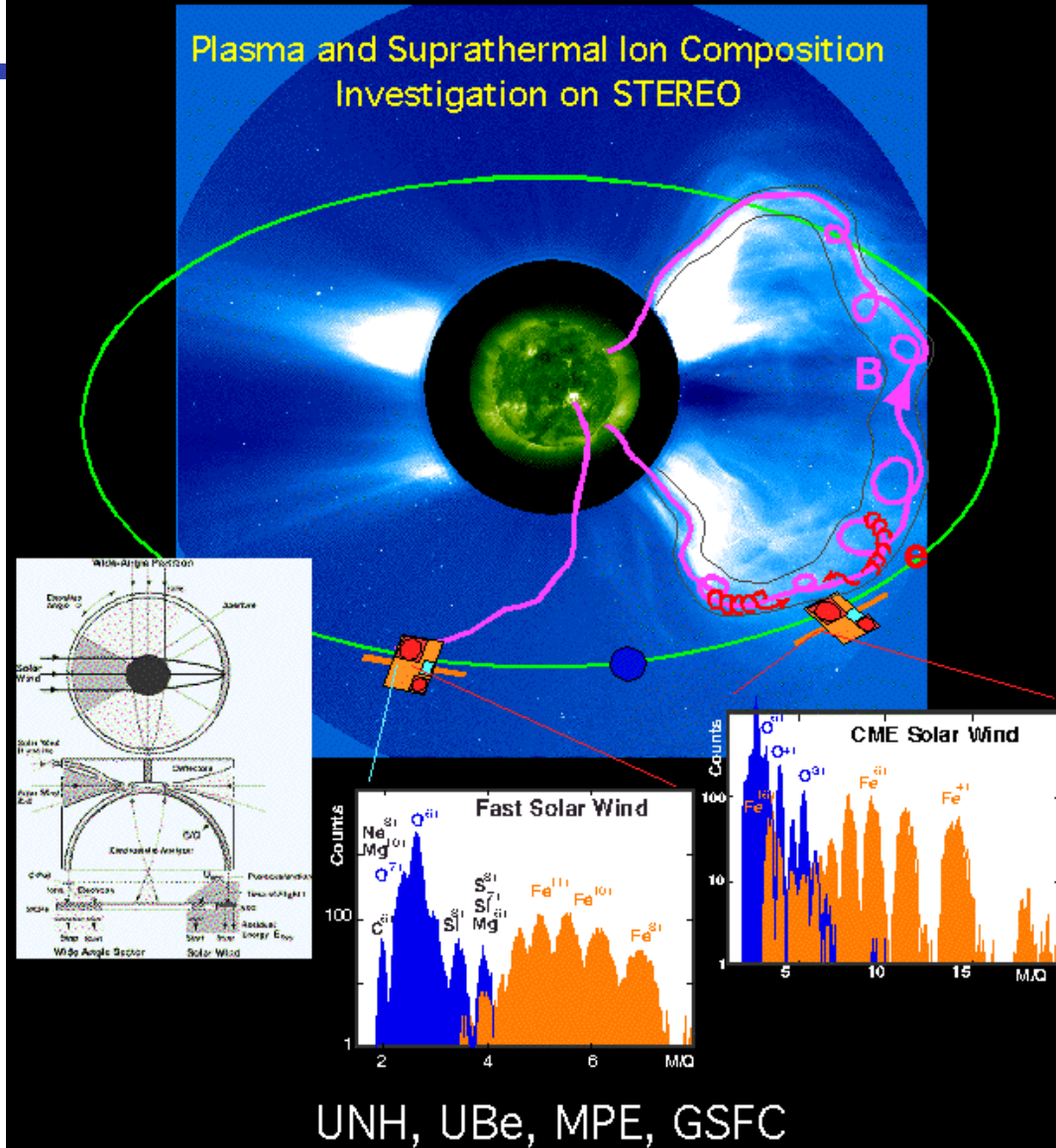
Wind Waves RAD2 receiver: 1999/6/29





# PLASTIC

## Plasma and Suprathermal Ion Composition Investigation on STEREO



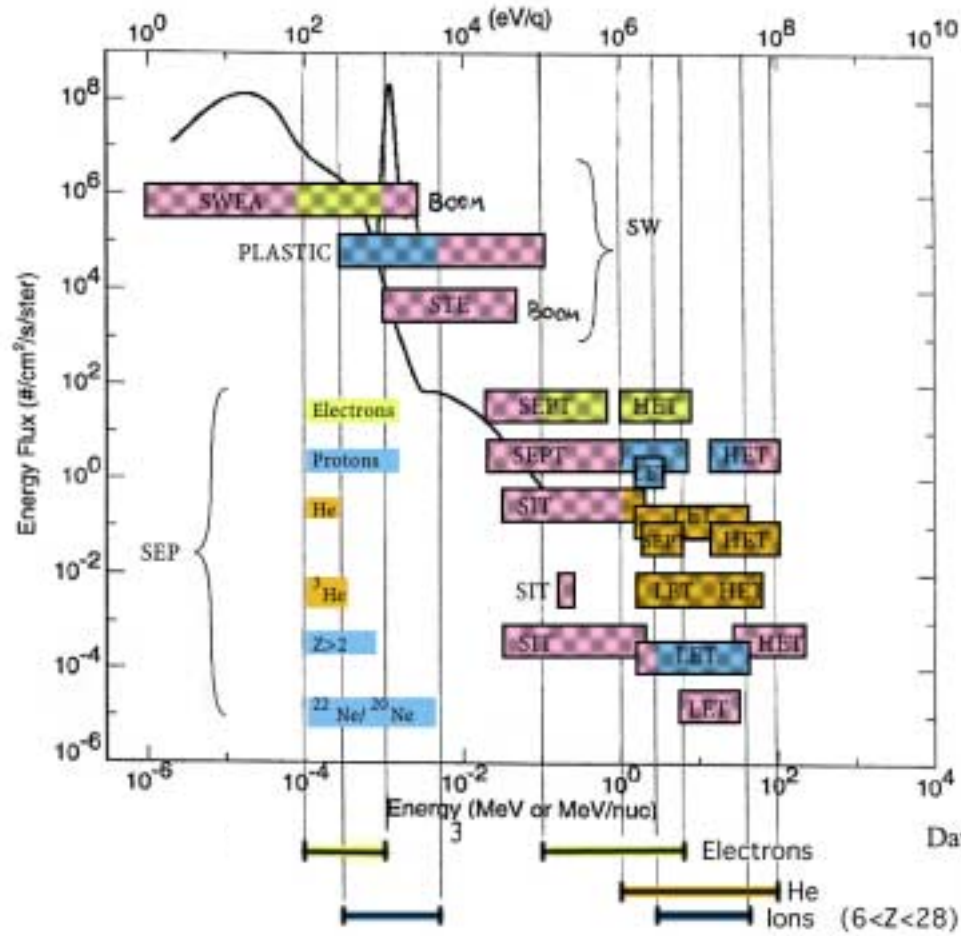




# STEREO IMPACT

System Requirements Review  
2000-May-24,25

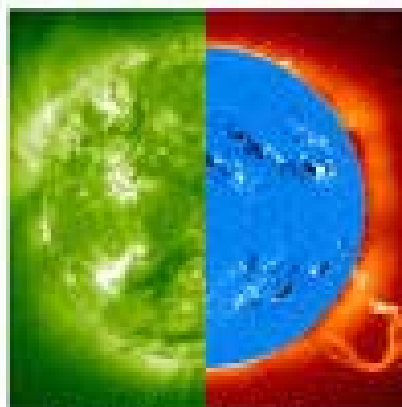
## IMPACT / PLASTIC Energy Coverage





## SECCHI Exploration of CMEs and the Heliosphere on STEREO

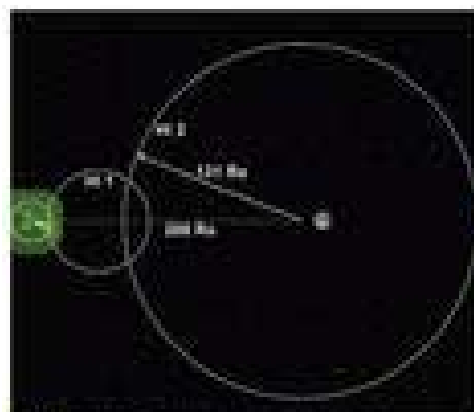
- What Configurations of the Corona Lead to a CME?
- What Initiates a CME?
- What Accelerates CMEs?
- How Does a CME Interact With the Heliosphere?
- How do CMEs Cause Space Weather Disturbances?



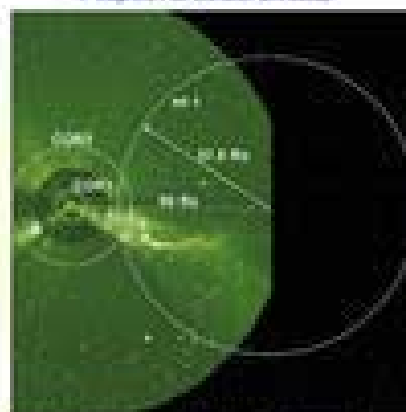
- Explore the Magnetic Origin of CMEs
  - Photospheric Streaming Motions
  - Magnetic Flux Emergence
  - Magnetic Flux Evolution and Decay



- Understand the Interaction of CMEs
  - Reconnection
  - The Role of Flares on Magnetic Field Effects
  - Speed vs. Mass Correlations



- The Sun-Earth Connection: Understand the Role of CMEs in Space Weather
  - Orbital Trajectory of Earth-Directed CMEs
  - Photo-Auroral Time and Case of Backscatter of CMEs

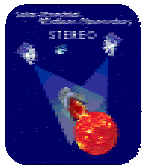


- Investigate the Interaction of CMEs With the Heliosphere
  - CME Physical Signatures at 1 AU
  - Generation of Shocks
  - Acceleration of Charged Particles
  - Interaction With Interplanetary Plasma Shear & Co-Rotating Interaction Regions
  - Interaction With Other CMEs



- Study the Physical Function of CMEs
  - Reconnection
  - Coupled Energy Input and Mass Injection
  - Effect on Heliospheric Structures

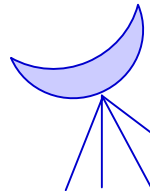
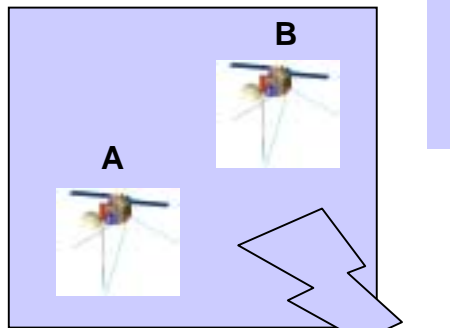




# STEREO as ILWS Test-bed?



NASA will provide two spacecraft that continuously transmit space weather measurements to a ground processing facility



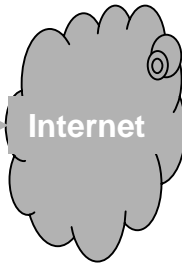
3 hours per day

DSN

STEREO  
Science  
Center  
NASA-GSFC



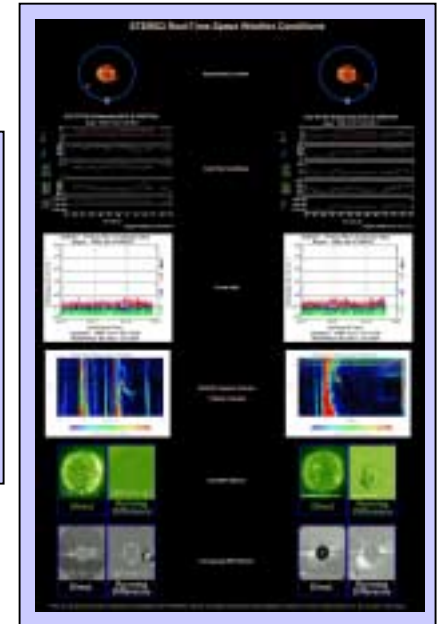
21 hours per day



Internet

Antenna  
Partners

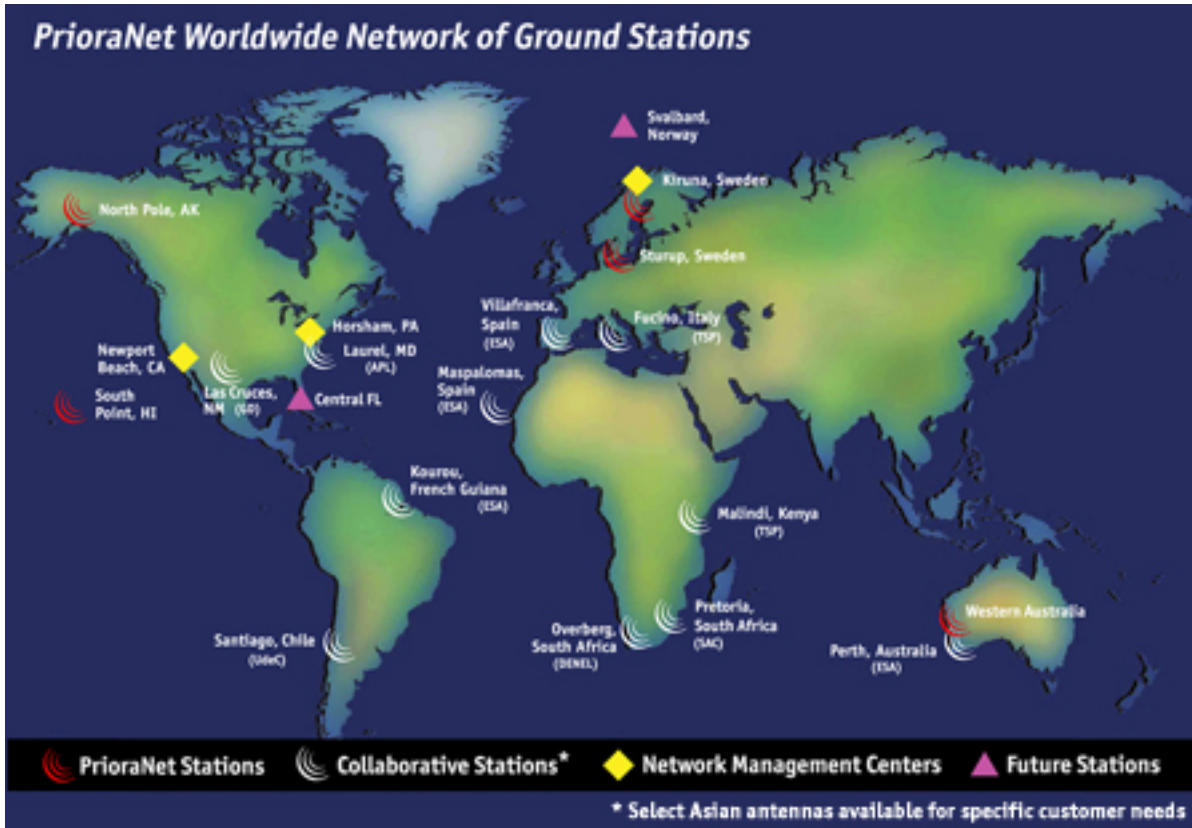
In Search Of:  
Antenna Partners



Real-time Space Weather  
Display and Archive







11m X-band antenna  
\$575 per hour

20 hours per day X 2  
 spacecraft =  
**\$8.4M/year**



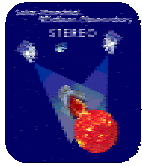


# Potential Antenna Partners



- Rutherford Appleton Labs (U.K.) ACE antenna
    - Upgrade to X-band underway
  - NOAA Space Environment Center
- 
- U.C. Berkeley (STEREO PI team)
    - RHESSI antenna requires upgrade to X-band
  - CRL (Japan)
    - Budget and technical feasibility studies are underway
  - Germany (GSOC request through University of Kiel)
  - Brazil (W. Gonzales)
  - ESOC/ESA has a network of 15m X-band antennas with significant spare capacity (estimated ~200 hours per week)





# What Do Antenna Partners Need?



- **X-band dish and receiver**
  - 7 meter dish covers 2-year nominal mission
  - 15 meter dish covers 5-year extended mission
- **Low data rate Internet connection during telemetry reception periods**
  - TCP/IP socket, email, ftp, etc.
- **Beacon description reprints available**
- **<ftp://stereo.gsfc.nasa.gov/pub/cstcyr/STEREO/>**





# Antenna Contact Duration

(latitude versus time)



For a single mid-latitude location on Earth, the two STEREO spacecraft will be seasonally out-of-phase

Example using RAL, U.K. (lat. +51°)

	STEREO-A	STEREO-B
March 2006	6 hours/day	15 hours/day
Sept 2006	15 hours/day	6 hours/day
March 2007	6 hours/day	15 hours/day

